

88. (a) We write the centripetal acceleration (which is the same for each, since they have identical mass) as $r\omega^2$ where ω is the unknown angular speed. Thus,

$$\frac{G(M)(M)}{(2r)^2} = \frac{GM^2}{4r^2} = Mr\omega^2$$

which gives $\omega = \frac{1}{2}\sqrt{MG/r^3} = 2.2 \times 10^{-7} \text{ rad/s}$.

- (b) To barely escape means to have total energy equal to zero (see discussion prior to Eq. 14-27). If m is the mass of the meteoroid, then

$$\frac{1}{2}mv^2 - \frac{GmM}{r} - \frac{GmM}{r} = 0 \quad \Rightarrow \quad v = \sqrt{\frac{4GM}{r}} = 8.9 \times 10^4 \text{ m/s} .$$