

67. We may approximate the planets and their motions as particles in circular orbits, and use Eq. 12-26

$$L = \sum_{i=1}^9 \ell_i = \sum_{i=1}^9 m_i r_i^2 \omega_i$$

to compute the total angular momentum. Since we assume the angular speed of each one is constant, we have (in rad/s)  $\omega_i = 2\pi/T_i$  where  $T_i$  is the time for that planet to go around the Sun (this and related information is found in Appendix C but there, the  $T_i$  are expressed in years and we'll need to convert with  $3.156 \times 10^7$  s/y, and the  $M_i$  are expressed as multiples of  $M_{\text{earth}}$  which we'll convert by multiplying by  $5.98 \times 10^{24}$  kg).

(a) Using SI units, we find (with  $i = 1$  designating Mercury)

$$\begin{aligned} L &= \sum_{i=1}^9 m_i r_i^2 \left( \frac{2\pi}{T_i} \right) \\ &= 2\pi \frac{3.34 \times 10^{23}}{7.61 \times 10^6} (57.9 \times 10^9)^2 + 2\pi \frac{4.87 \times 10^{24}}{19.4 \times 10^7} (108 \times 10^9)^2 + \\ &\quad 2\pi \frac{5.98 \times 10^{24}}{3.156 \times 10^7} (150 \times 10^9)^2 + 2\pi \frac{6.40 \times 10^{23}}{5.93 \times 10^7} (228 \times 10^9)^2 + \\ &\quad 2\pi \frac{1.9 \times 10^{27}}{3.76 \times 10^8} (778 \times 10^9)^2 + 2\pi \frac{5.69 \times 10^{26}}{9.31 \times 10^8} (1430 \times 10^9)^2 + \\ &\quad 2\pi \frac{8.67 \times 10^{25}}{2.65 \times 10^9} (2870 \times 10^9)^2 + 2\pi \frac{1.03 \times 10^{26}}{5.21 \times 10^9} (4500 \times 10^9)^2 + \\ &\quad 2\pi \frac{1.2 \times 10^{22}}{7.83 \times 10^9} (5900 \times 10^9)^2 \\ &= 3.14 \times 10^{43} \text{ kg}\cdot\text{m}^2/\text{s} . \end{aligned}$$

(b) The fractional contribution of Jupiter is

$$\frac{\ell_5}{L} = \frac{2\pi \frac{1.9 \times 10^{27}}{3.76 \times 10^8} (778 \times 10^9)^2}{3.14 \times 10^{43}} = 0.61 .$$