

## Chapter 34 Even Answers

2.  $2.25 \times 10^8 \text{ m/s}$
4. 38.0 pT
6.  $E = (300 \text{ V/m}) \cos(62.8x - 1.88 \times 10^{10}t)$ ,  $B = (1.00 \text{ } \mu\text{T}) \cos(62.8x - 1.88 \times 10^{10}t)$
10.  $2.9 \times 10^8 \text{ m/s} \pm 5\%$
12. 49.5 mV in amplitude
14. (a) 13.3 nJ/m<sup>3</sup> (b) 13.3 nJ/m<sup>3</sup> (c) 7.96 W/m<sup>2</sup>
16.  $5.16 \times 10^{-10} \text{ T}$ ,  $\sim 10^5$  times weaker than the Earth's field.
18. (a)  $\mathbf{E} \cdot \mathbf{B} = 0$  (b)  $(11.5 \mathbf{i} - 28.6 \mathbf{j}) \text{ W/m}^2$
20. (a) 50.0% (b) 269 kW/m<sup>2</sup> toward the oven chamber  
(c) 14.2 kV/m
22. 5.16 m
24. (a) 540 V/m (b)  $2.58 \text{ } \mu\text{J/m}^3$   
(c) 773 W/m<sup>2</sup> (d) 77.3% of the flux in Example 34.5
26. (a)  $1.60 \times 10^{-10} \text{ kg} \cdot \text{m/s}$  (b)  $1.60 \times 10^{-10} \text{ N}$
28.  $6.67 \times 10^{-10} \text{ N}$
30. (a) 577 W/m<sup>2</sup> (b)  $2.06 \times 10^{16} \text{ W}$   
(c)  $6.87 \times 10^7 \text{ N}$  if Mars behaves as a perfect absorber.  
(d) The gravitational force is  $\sim 10^{13}$  times stronger than the light force, in the opposite direction.
32. 3.48 A/m<sup>2</sup>
34. (a) 93.3% (b) 50.0% (c) 0
36. (a) 134 m (b) 46.9 m
38.  $\frac{2\pi mc}{qB}$
40. (a)  $\sim 10^8 \text{ Hz}$ , radio wave (b)  $\sim 10^{13} \text{ Hz}$ , infrared

42. (a) 0.690 wavelengths (b) 58.9 wavelengths.
44. 60.0 km
46.  $1.00 \times 10^3$  km (621 mi), not very practical.
48. (a) 4.17 m to 4.55 m (b) 3.41 m to 3.66 m (c) 1.61 m to 1.67 m
50.  $\sim 10^6$  J
52. (a) See solution (b)  $3.78 \times 10^{-7}$  m (or 378 nm)
54. (a)  $\frac{(\Delta V)i}{2\pi r l}$  radially outward (b)  $(\Delta V)i$  (c) The Poynting vector is now directed radially inward.
56. (a)  $3.14 \times 10^7$  W (b)  $0.625$  W/m<sup>2</sup> (c) 0.513%
58. (a)  $23.9$  W/m<sup>2</sup> (b) It is 4.19 times the standard.
60. (a) 388 K (b) 363 K
62. (a)  $6.16 \times 10^{-6}$  Pa (b)  $1.64 \times 10^{10}$  times smaller than atmospheric pressure
64. (a)  $\frac{4\rho gc}{3} \left(\frac{3m}{4\pi\rho}\right)^{1/3}$  (b)  $\frac{4\pi r^2 \rho gc}{3} \left(\frac{3m}{4\pi\rho}\right)^{1/3}$
66. (a) 1.50 cm (b) 25.0  $\mu$ J (c) 7.37 mJ/m<sup>3</sup>  
(d) 40.8 kV/m, 136  $\mu$ T (e) 83.3  $\mu$ N
68.  $6.37 \times 10^{-7}$  Pa
70. The projected area is  $\pi r^2$  and the radiating area is  $4\pi r^2$ .  
Orbital radius,  $R = 4.77 \times 10^9$  m = 4.77 Gm
72. (a) 3.33 m, 11.1 ns, 6.67 pT  
(b)  $\mathbf{E} = (2.00 \text{ mV/m}) \cos 2\pi \left( \frac{x}{3.33 \text{ m}} - \frac{t}{11.1 \text{ ns}} \right) \mathbf{j}$ ,  $\mathbf{B} = (6.67 \text{ pT}) \cos 2\pi \left( \frac{x}{3.33 \text{ m}} - \frac{t}{11.1 \text{ ns}} \right) \mathbf{k}$   
(c)  $5.31 \times 10^{-9}$  W/m<sup>2</sup> (d)  $1.77 \times 10^{-17}$  J/m<sup>3</sup> (e)  $3.54 \times 10^{-17}$  Pa