

CHAPTER 12.

1. With a potential of the form

$$V(r) = \frac{1}{2} m \omega^2 r^2$$

the perturbation reduces to

$$\begin{aligned} H_1 &= \frac{1}{2m^2c^2} \mathbf{S} \cdot \mathbf{L} \frac{1}{r} \frac{dV(r)}{dr} = \frac{\omega^2}{4mc^2} (\mathbf{J}^2 - \mathbf{L}^2 - \mathbf{S}^2) \\ &= \frac{(\hbar\omega)^2}{4mc^2} (j(j+1) - l(l+1) - s(s+1)) \end{aligned}$$

where l is the orbital angular momentum, s is the spin of the particle in the well (e.g. $1/2$ for an electron or a nucleon) and j is the total angular momentum. The possible values of j are $l + s, l + s - 1, l + s - 2, \dots, |l - s|$.

The unperturbed energy spectrum is given by $E_{n_r, l} = \hbar\omega(2n_r + l + \frac{3}{2})$. Each of the levels characterized by l is $(2l + 1)$ -fold degenerate, but there is an additional degeneracy, not unlike that appearing in hydrogen. For example $n_r=2, l=0$, $n_r=1, l=2$, $n_r=0, l=4$ all have the same energy.

A picture of the levels and their spin-orbit splitting is given below.

2. The effects that enter into the energy levels corresponding to $n = 2$, are (I) the basic Coulomb interaction, (ii) relativistic and spin-orbit effects, and (iii) the hyperfine structure which we are instructed to ignore. Thus, in the absence of a magnetic field, the levels under the influence of the Coulomb potential consist of $2n^2 = 8$ degenerate levels. Two of the levels are associated with $l = 0$ (spin up and spin down) and six

levels with $l = 0$, corresponding to $m_l = 1, 0, -1$, spin up and spin down. The latter can be rearranged into states characterized by \mathbf{J}^2 , \mathbf{L}^2 and J_z . There are two levels characterized by $j = l - 1/2 = 1/2$ and four levels with $j = l + 1/2 = 3/2$. These energies are split by relativistic effects and spin-orbit coupling, as given in Eq. (12-16). We ignore reduced mass effects (other than in the original Coulomb energies). We therefore have

$$\begin{aligned}\Delta E &= -\frac{1}{2} m_e c^2 \alpha^4 \frac{1}{n^3} \left(\frac{1}{j+1/2} - \frac{3}{4n} \right) \\ &= -\frac{1}{2} m_e c^2 \alpha^4 \left(\frac{5}{64} \right) \quad j = 1/2 \\ &= -\frac{1}{2} m_e c^2 \alpha^4 \left(\frac{1}{64} \right) \quad j = 3/2\end{aligned}$$

(b) The Zeeman splittings for a given j are

$$\begin{aligned}\Delta E_B &= \frac{e\hbar B}{2m_e} m_j \left(\frac{2}{3} \right) \quad j = 1/2 \\ &= \frac{e\hbar B}{2m_e} m_j \left(\frac{4}{3} \right) \quad j = 3/2\end{aligned}$$

Numerically $\frac{1}{128} m_e c^2 \alpha^4 \approx 1.132 \times 10^{-5} \text{ eV}$, while for $B = 2.5 \text{ T}$ $\frac{e\hbar B}{2m_e} = 14.47 \times 10^{-5} \text{ eV}$,

so under these circumstances the magnetic effects are a factor of 13 larger than the relativistic effects. Under these circumstances one could neglect these and use Eq. (12-26).

3. The unperturbed Hamiltonian is given by Eq. (12-34) and the magnetic field interacts both with the spin of the electron and the spin of the proton. This leads to

$$H = A \frac{\mathbf{S} \cdot \mathbf{I}}{\hbar^2} + a \frac{S_z}{\hbar} + b \frac{I_z}{\hbar}$$

Here

$$\begin{aligned}A &= \frac{4}{3} \alpha^4 m_e c^2 g_p \left(\frac{m_e}{M_p} \right) \frac{\mathbf{I} \cdot \mathbf{S}}{\hbar^2} \\ a &= 2 \frac{e\hbar B}{2m_e} \\ b &= -g_p \frac{e\hbar B}{2M_p}\end{aligned}$$

Let us now introduce the total spin $\mathbf{F} = \mathbf{S} + \mathbf{I}$. It follows that

$$\begin{aligned}\frac{\mathbf{S} \cdot \mathbf{I}}{\hbar^2} &= \frac{1}{2\hbar^2} \left(\hbar^2 F(F+1) - \frac{3}{4} \hbar^2 - \frac{3}{4} \hbar^2 \right) \\ &= \frac{1}{4} \quad \text{for } F=1 \\ &= -\frac{3}{4} \quad \text{for } F=0\end{aligned}$$

We next need to calculate the matrix elements of $aS_z + bI_z$ for eigenstates of \mathbf{F}^2 and F_z . These will be exactly like the spin triplet and spin singlet eigenstates. These are

$$\begin{aligned}\langle 1,1 | aS_z + bI_z | 1,1 \rangle &= \langle \chi_+ \xi_+ | aS_z + bI_z | \chi_+ \xi_+ \rangle = \frac{1}{2} (a+b) \\ \langle 1,0 | aS_z + bI_z | 1,0 \rangle &= \left(\frac{1}{\sqrt{2}} \right)^2 \langle \chi_+ \xi_- + \chi_- \xi_+ | aS_z + bI_z | \chi_+ \xi_- + \chi_- \xi_+ \rangle = 0 \\ \langle 1,-1 | aS_z + bI_z | 1,-1 \rangle &= \langle \chi_- \xi_- | aS_z + bI_z | \chi_- \xi_- \rangle = -\frac{1}{2} (a+b)\end{aligned}$$

And for the singlet state ($F=0$)

$$\begin{aligned}\langle 1,0 | aS_z + bI_z | 0,0 \rangle &= \left(\frac{1}{\sqrt{2}} \right)^2 \langle \chi_+ \xi_- + \chi_- \xi_+ | aS_z + bI_z | \chi_+ \xi_- - \chi_- \xi_+ \rangle = \frac{1}{2} (a-b) \\ \langle 0,0 | aS_z + bI_z | 0,0 \rangle &= \left(\frac{1}{\sqrt{2}} \right)^2 \langle \chi_+ \xi_- - \chi_- \xi_+ | aS_z + bI_z | \chi_+ \xi_- - \chi_- \xi_+ \rangle = 0\end{aligned}$$

Thus the magnetic field introduces mixing between the $|1,0\rangle$ state and the $|0,0\rangle$ state. We must therefore diagonalize the submatrix

$$\begin{pmatrix} A/4 & (a-b)/2 \\ (a-b)/2 & -3A/4 \end{pmatrix} = \begin{pmatrix} -A/4 & 0 \\ 0 & -A/4 \end{pmatrix} + \begin{pmatrix} A/2 & (a-b)/2 \\ (a-b)/2 & -A/2 \end{pmatrix}$$

The second submatrix commutes with the first one. Its eigenvalues are easily determined to be $\pm \sqrt{A^2/4 + (a-b)^2/4}$ so that the overall eigenvalues are

$$-A/4 \pm \sqrt{A^2/4 + (a-b)^2/4}$$

Thus the spectrum consists of the following states:

$$F = 1, F_z = 1 \quad E = A / 4 + (a + b) / 2$$

$$F = 1, F_z = -1 \quad E = A / 4 - (a + b) / 2$$

$$F = 1, 0; F_z = 0 \quad E = -A / 4 \pm \sqrt{(A^2 / 4 + (a - b)^2 / 4}$$

We can now put in numbers.

For $B = 10^{-4}$ T, the values, in units of 10^{-6} eV are 1.451, 1.439, $0(10^{-10})$, -2.89

For $B = 1$ T, the values in units of 10^{-6} eV are 57.21, -54.32, 54.29 and 7×10^{-6} .

4. According to Eq. (12-17) the energies of hydrogen-like states, including relativistic + spin-orbit contributions is given by

$$E_{n,j} = -\frac{1}{2} \frac{m_e c^2 (Z\alpha)^2}{(1 + m_e / M_p)} \frac{1}{n^2} - \frac{1}{2} m_e c^2 (Z\alpha)^4 \frac{1}{n^3} \left(\frac{1}{j + 1/2} - \frac{3}{4n} \right)$$

The wavelength in a transition between two states is given by

$$\lambda = \frac{2\pi\hbar c}{\Delta E}$$

where ΔE is the change in energy in the transition. We now consider the transitions $n=3, j=3/2 \rightarrow n=1, j=1/2$ and $n=3, j=1/2 \rightarrow n=1, j=1/2$. The corresponding energy differences (neglecting the reduced mass effect) is

$$(3, 3/2 \rightarrow 1, 1/2) \quad \Delta E = \frac{1}{2} m_e c^2 (Z\alpha)^2 \left(1 - \frac{1}{9}\right) + \frac{1}{2} m_e c^2 (Z\alpha)^4 \frac{1}{4} \left(1 - \frac{1}{27}\right)$$

$$(3, 1/2 \rightarrow 1, 1/2) \quad \frac{1}{2} m_e c^2 (Z\alpha)^2 \left(1 - \frac{1}{9}\right) + \frac{1}{2} m_e c^2 (Z\alpha)^4 \frac{1}{4} \left(1 - \frac{3}{27}\right)$$

We can write these in the form

$$(3, 3/2 \rightarrow 1, 1/2) \quad \Delta E_0 \left(1 + \frac{13}{48} (Z\alpha)^2\right)$$

$$(3, 1/2 \rightarrow 1, 1/2) \quad \Delta E_0 \left(1 + \frac{1}{18} (Z\alpha)^2\right)$$

where

$$\Delta E_0 = \frac{1}{2} m_e c^2 (Z\alpha)^2 \frac{8}{9}$$

The corresponding wavelengths are

$$(3,3/2 \rightarrow 1,1/2) \quad \lambda_0(1 - \frac{13}{48}(Z\alpha)^2) = 588.995 \times 10^{-9} m$$

$$(3,1/2 \rightarrow 1,1/2) \quad \lambda_0(1 - \frac{1}{18}(Z\alpha)^2) = 589.592 \times 10^{-9} m$$

We may use the two equations to calculate λ_0 and Z . Dividing one equation by the other we get, after a little arithmetic $Z = 11.5$, which fits with the $Z = 11$ for Sodium.

(Note that if we take for λ_0 the average of the two wavelengths, then, using $\lambda_0 = 2\pi\hbar c / \Delta E_0 = 9\pi\hbar / 2mc(Z\alpha)^2$, we get a seemingly unreasonably small value of $Z = 0.4$! This is not surprising. The ionization potential for sodium is 5.1 eV instead of $Z^2(13.6 \text{ eV})$, for reasons that will be discussed in Chapter 14)

4. The relativistic correction to the kinetic energy term is $-\frac{1}{2mc^2} \left(\frac{\mathbf{p}^2}{2m} \right)^2$. The energy shift in the ground state is therefore

$$\Delta E = -\frac{1}{2mc^2} \langle 0 | \left(\frac{\mathbf{p}^2}{2m} \right)^2 | 0 \rangle = -\frac{1}{2mc^2} \langle 0 | (H - \frac{1}{2}m\omega^2 r^2)^2 | 0 \rangle$$

To calculate $\langle 0 | r^2 | 0 \rangle$ and $\langle 0 | r^4 | 0 \rangle$ we need the ground state wave function. We know that for the one-dimensional oscillator it is

$$u_0(x) = \left(\frac{m\omega}{\pi\hbar} \right)^{1/4} e^{-m\omega x^2 / 2\hbar}$$

so that for the three dimensional oscillator it is

$$u_0(r) = u_0(x)u_0(y)u_0(z) = \left(\frac{m\omega}{\pi\hbar} \right)^{3/4} e^{-m\omega r^2 / 2\hbar}$$

It follows that

$$\begin{aligned} \langle 0 | r^2 | 0 \rangle &= \int_0^\infty 4\pi r^2 dr \left(\frac{m\omega}{\pi\hbar} \right)^{3/2} r^2 e^{-m\omega r^2 / \hbar} = \\ &= 4\pi \left(\frac{m\omega}{\pi\hbar} \right)^{3/2} \left(\frac{\hbar}{m\omega} \right)^{5/2} \int_0^\infty dy y^4 e^{-y^2} = \\ &= \frac{3\hbar}{2m\omega} \end{aligned}$$

We can also calculate

$$\begin{aligned}
 \langle 0 | r^4 | 0 \rangle &= \int_0^\infty 4\pi r^2 dr \left(\frac{m\omega}{\pi\hbar} \right)^{3/2} r^4 e^{-m\omega r^2/\hbar} = \\
 &= 4\pi \left(\frac{m\omega}{\pi\hbar} \right)^{3/2} \left(\frac{\hbar}{m\omega} \right)^{7/2} \int_0^\infty dy y^6 e^{-y^2} \\
 &= \frac{15}{4} \left(\frac{\hbar}{m\omega} \right)^2
 \end{aligned}$$

We made use of $\int_0^\infty dz z^n e^{-z} = \Gamma(n+1) = n\Gamma(n)$ and $\Gamma(\frac{1}{2}) = \sqrt{\pi}$

Thus

$$\begin{aligned}
 \Delta E &= -\frac{1}{2mc^2} \left(\left(\frac{3}{2} \hbar\omega \right)^2 - \left(\frac{3}{2} \hbar\omega \right) \left(m\omega^2 \frac{3\hbar}{2m\omega} \right) + \frac{1}{4} m^2 \omega^4 \left(\frac{15}{4} \right) \left(\frac{\hbar}{m\omega} \right)^2 \right) \\
 &= -\frac{15}{32} \frac{(\hbar\omega)^2}{mc^2}
 \end{aligned}$$

6. (a) With $J = 1$ and $S = 1$, the possible values of the orbital angular momentum, such that $j = L + S, L + S - 1, \dots, |L - S|$ can only be $L = 0, 1, 2$. Thus the possible states are $^3S_1, ^3P_1, ^3D_1$. The parity of the deuteron is $(-1)^L$ assuming that the intrinsic parities of the proton and neutron are taken to be $+1$. Thus the S and D states have positive parity and the P state has opposite parity. Given parity conservation, the only possible admixture can be the 3D_1 state.

(b) The interaction with a magnetic field consists of three contributions: the interaction of the spins of the proton and neutron with the magnetic field, and the $\mathbf{L} \cdot \mathbf{B}$ term, if L is not zero. We write

$$H = -\mathbf{M}_p \cdot \mathbf{B} - \mathbf{M}_n \cdot \mathbf{B} - \mathbf{M}_L \cdot \mathbf{B}$$

$$\mathbf{M}_p = \frac{eg_p}{2M} \mathbf{S}_p = (5.5792) \frac{e\hbar}{2M} \left(\frac{\mathbf{S}_p}{\hbar} \right)$$

$$\text{where } \mathbf{M}_n = \frac{eg_n}{2M} \mathbf{S}_n = (-3.8206) \frac{e\hbar}{2M} \left(\frac{\mathbf{S}_n}{\hbar} \right)$$

$$\mathbf{M}_L = \frac{e}{2M_{red}} \mathbf{L}$$

We take the neutron and proton masses equal ($= M$) and the reduced mass of the two-particle system for equal masses is $M/2$. For the 3S_1 state, the last term does not contribute.

If we choose \mathbf{B} to define the z axis, then the energy shift is

$$-\frac{eB\hbar}{2M} \langle {}^3S_1 | g_p \left(\frac{S_{pz}}{\hbar} \right) + g_n \left(\frac{S_{nz}}{\hbar} \right) | {}^3S_1 \rangle$$

We write

$$g_p \left(\frac{S_{pz}}{\hbar} \right) + g_n \left(\frac{S_{nz}}{\hbar} \right) = \frac{g_p + g_n}{2} \frac{S_{pz} + S_{nz}}{\hbar} + \frac{g_p - g_n}{2} \frac{S_{pz} - S_{nz}}{\hbar}$$

It is easy to check that the last term has zero matrix elements in the triplet states, so that we are left with $\frac{1}{2}(g_p + g_n) \frac{S_z}{\hbar}$, where S_z is the z -component of the total spin..

Hence

$$\langle {}^3S_1 | H_1 | {}^3S_1 \rangle = -\frac{3B\hbar}{2M} \frac{g_p + g_n}{2} m_s$$

where m_s is the magnetic quantum number ($m_s = 1, 0, -1$) for the total spin. We may therefore write the magnetic moment of the deuteron as

$$\mu_{eff} = -\frac{e}{2M} \frac{g_p + g_n}{2} \mathbf{S} = -(0.8793) \frac{e}{2M} \mathbf{S}$$

The experimental measurements correspond to $g_d = 0.8574$ which suggests a small admixture of the 3D_1 to the deuteron wave function.